

EUV MULTILAYER COATINGS FOR SPACE APPLICATIONS

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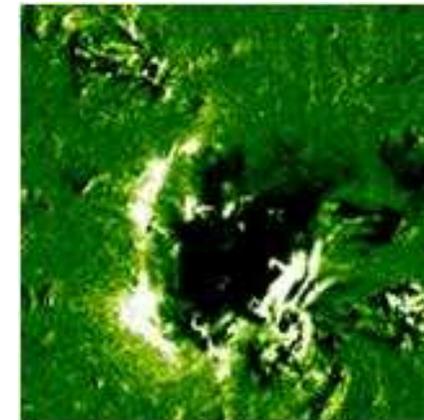
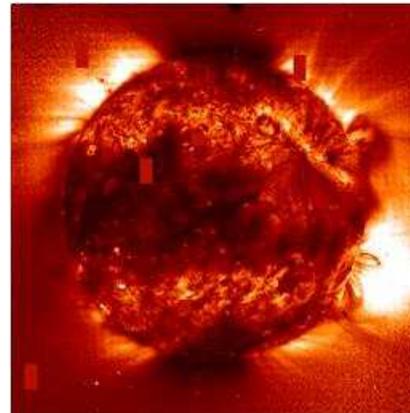
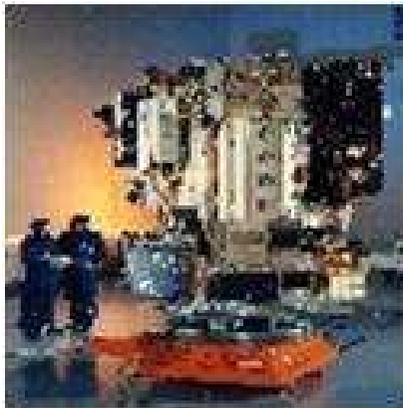
EUV multilayer coatings applications

special and extreme

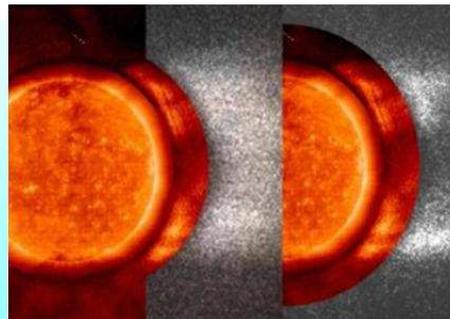
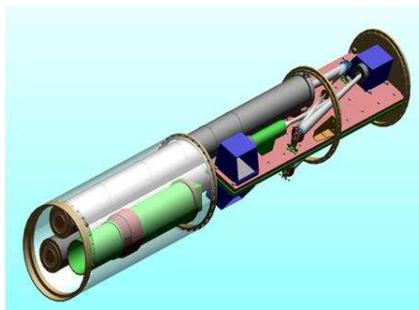
EUVL
HHG
SPACE
FEL

Application to solar physics

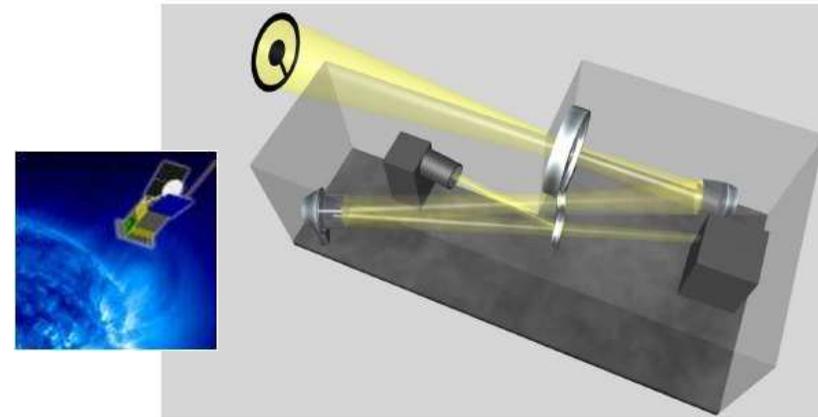
PAST: SOHO EIT, TRACE



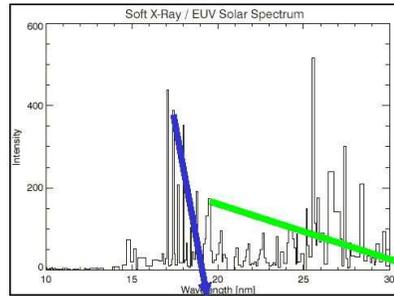
NOW: SCORE
SOUNDING ROCKET



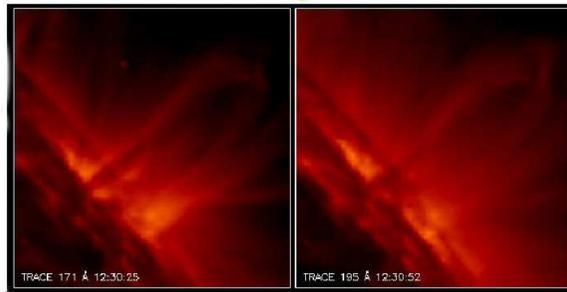
FUTURE: SOLO METIS



Application to Solar physics



Fe IX,
 $\lambda=171\text{\AA}$
 $\log T=5.9$



Fe XII,
 $\lambda=195\text{\AA}$
 $\log T=6.2$

Solar telescopes need to be tuned to specific coronal emission lines to allow “monochromatic” high-resolution imaging.

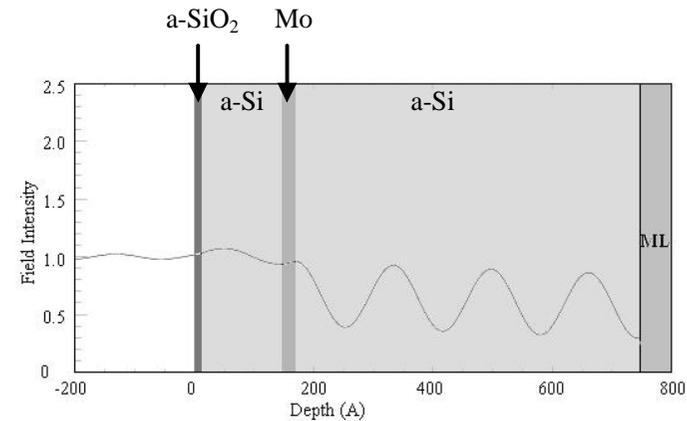
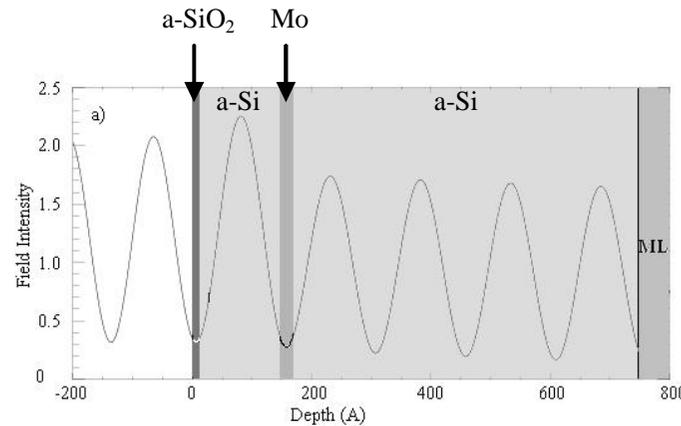
- Multilayer for solar line observations have been fabricated at Fe-IX (17.1 nm), Fe-XII (19.5 nm), Fe-XV (28.4 nm) and He-II (30.4 nm).
- Performance of such MLs have been evaluated in terms of **peak reflectivity** at working wavelength and **rejection capability** of unwanted lines
- Capping layer can be also important for lifetime

Application to Solar physics

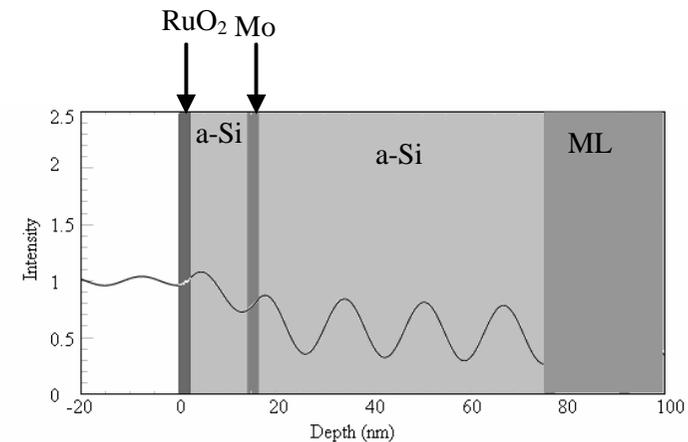
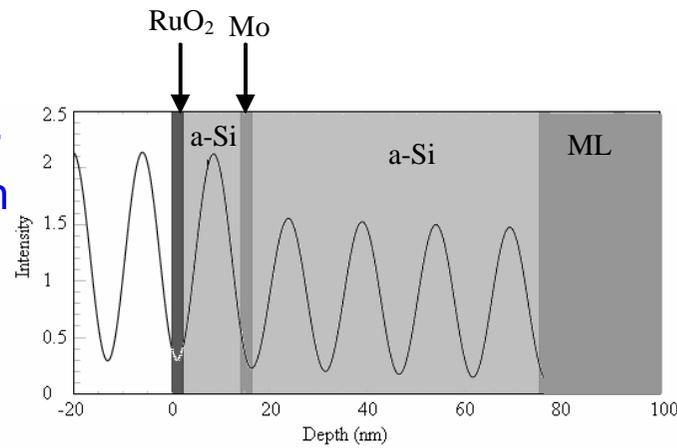
$\lambda_{\text{noise}} \rightarrow 30.4 \text{ nm (He-II)}$

$\lambda_{\text{peak}} \rightarrow 28.4 \text{ nm (Fe-XV)}$

Mo/Si CL

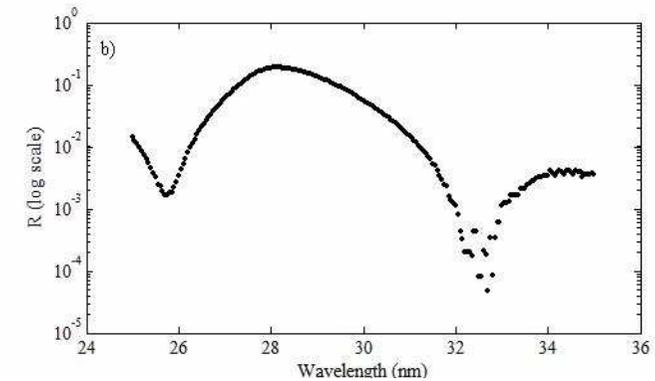
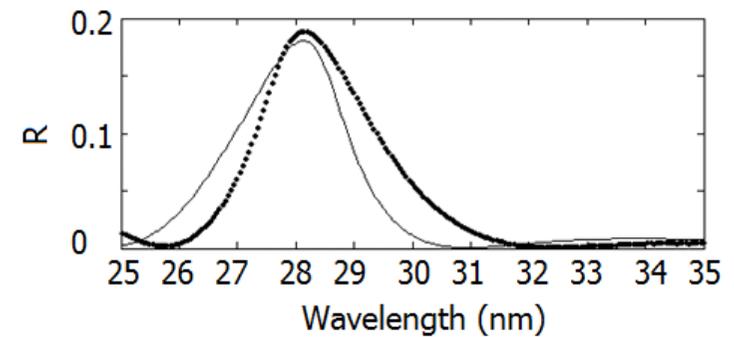
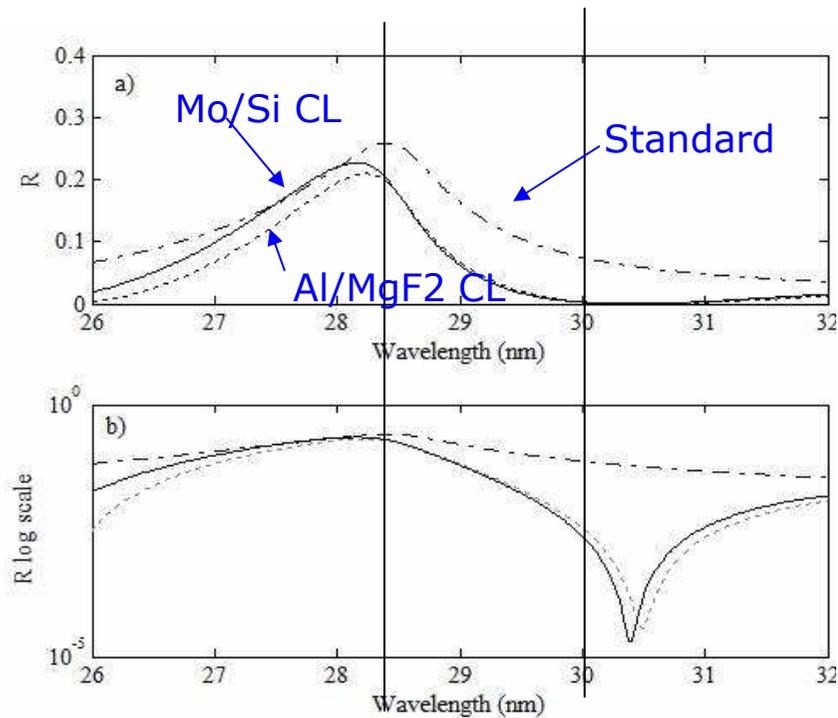


Mo/Si CL + Ru layer
on top for protection



Application to Solar physics

Mo/Si CL



Application to Solar physics

SCORE sounding rocket experiment

ML for multiband observations (30.4, 121.6, 450-600)

SiC/Mg
MLS

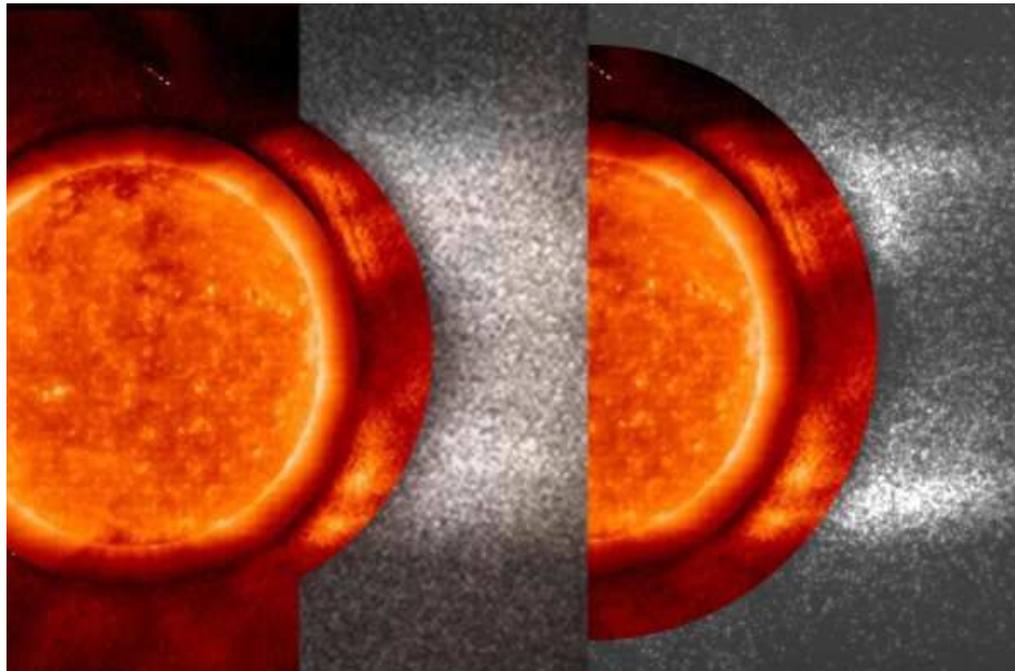


Application to Solar physics

SCORE sounding rocket experiment

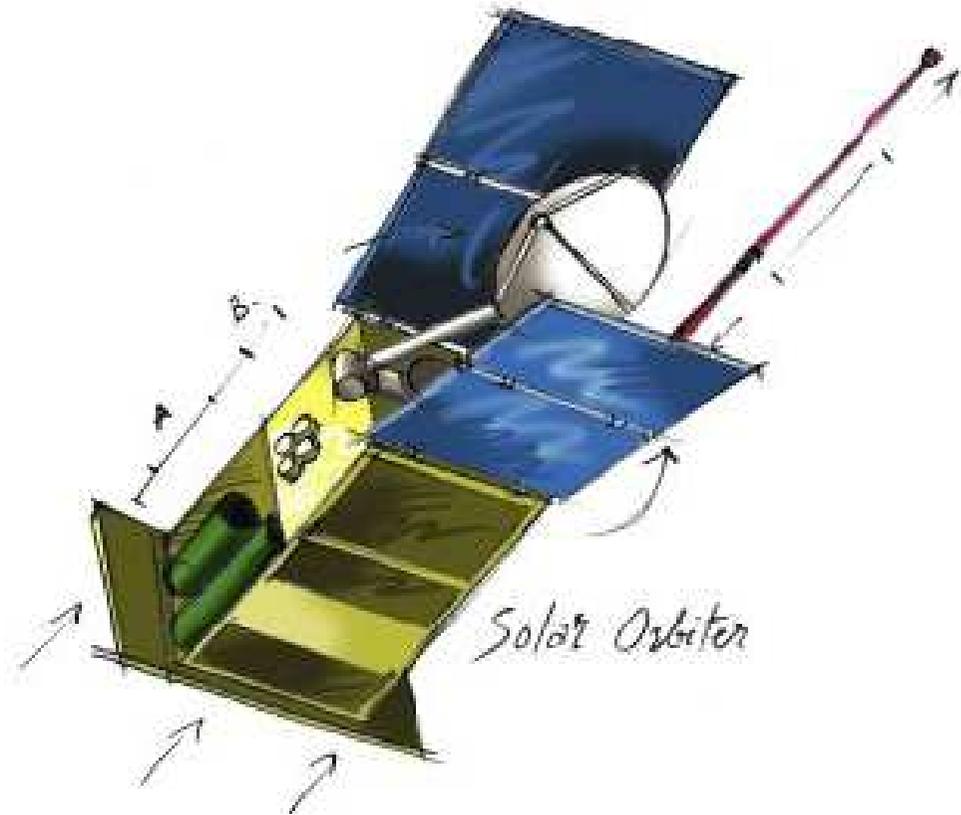
**cromosphere (HEIT), inner
corona (HECOR) @ He 304
A and outer corona
(SCORE) @ HI 1216 A**

**cromosphere (HEIT), inner corona
(HECOR) @ He 304 A and outer
corona (SCORE) @ He 304 A**



SOLar Orbiter (SOLO)

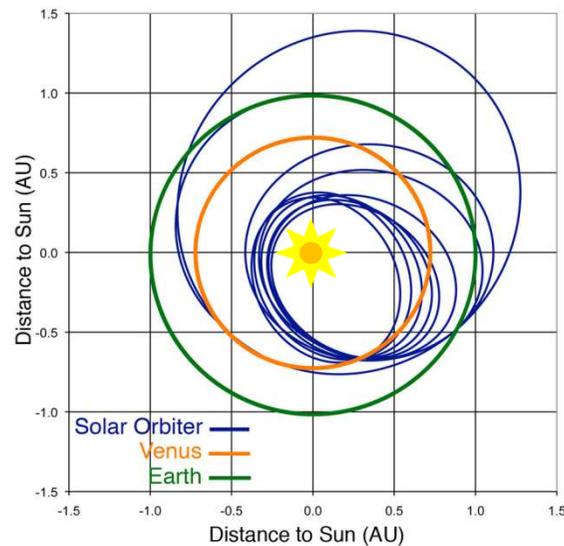
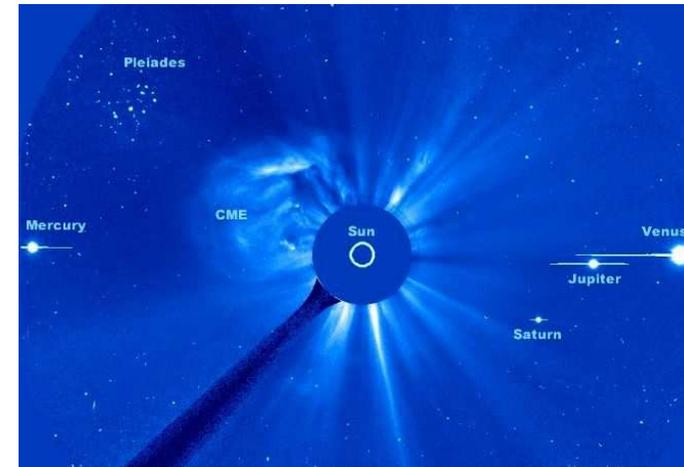
SOLar Orbiter mission is candidate for launch in January 2017. Its orbiting around the Sun will last about 7,5 years and will take it as close as 0,23 A.U, with increasing inclination up to more than 30° with respect to the solar equator, being able to view the solar atmosphere with high spatial resolution, covering the poles and the sun-side not visible from Earth.



SOLO structure drawing

SOLO environment operation

- SOLO will be the closest mission to the sun
- Minimum perihelium will be 0.23 AU
- Temperature of MO: 236°
- Ion bombardment, H+ energy: 1 KeV

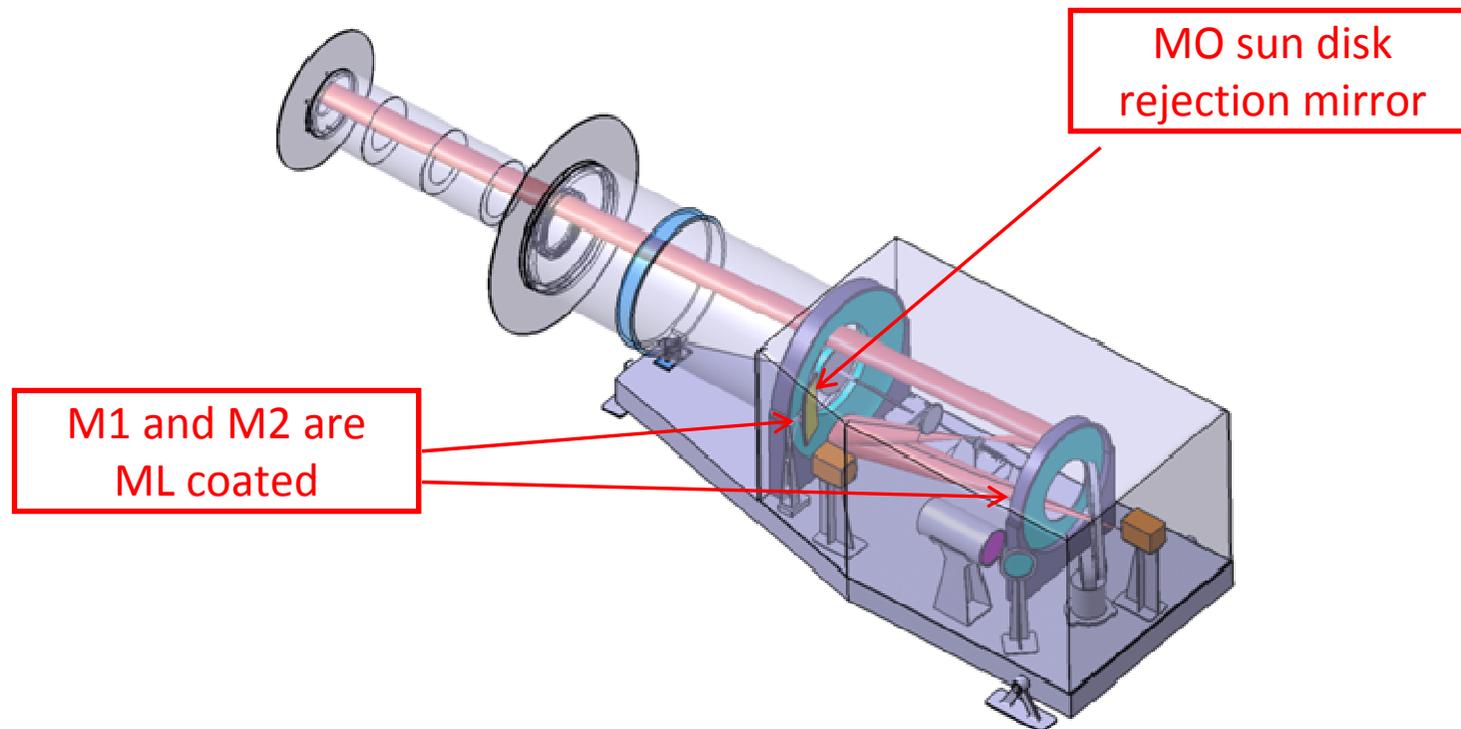


ESA Solar Orbiter Enviromental Specification, Issue2, 2008

	At 1AU (in the earth)	average
H+ Density (cm ⁻³)	8,7	25
H+ Speed (km/s)	468	468
N _{alpha} /N _{proton}	0,047	0,047
N _{O⁶⁺} /N _{proton}	0,0003	0,0003
N _{Fe¹⁰⁺} /N _{proton}	8,77 x 10 ⁻⁶	8,77 x 10 ⁻⁶

METIS on board of SOLar Orbiter

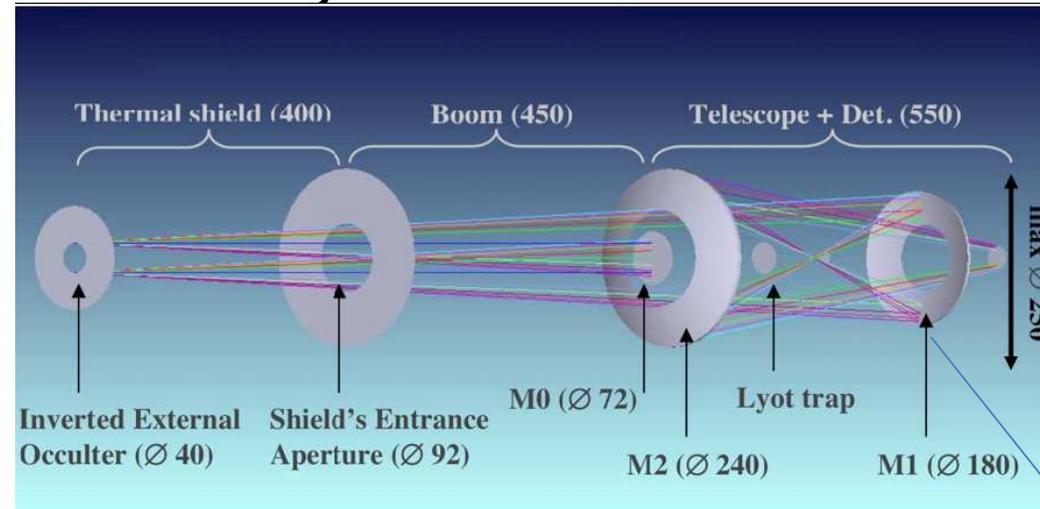
The Multi Element Telescope for Imaging and Spectroscopy (METIS) is an inverted – occultation coronagraph on board of the SOLO payload. It is devoted to image the solar corona in three different spectral ranges at 30,4 nm (He-II Lyman – α line), 121,6 nm (H-I Lyman – α line) and in the visible range (450 – 600 nm) by using a combination of band pass filter and high – reflectivity multilayer mirrors.



METIS drawing

Multilayers for METIS

METIS optical drawing



The telescope design consists of two concave mirrors with ring shape. M0, M1 and M2 have to be multilayer coated.

Requirements:

- High peak reflectivity at 30,4 nm, 121,6 nm, and in the visible range (450 – 600 nm)
- Stability over time
- Thermal stability
- High resistance to ion bombardments



M1 mirror Nicolosi

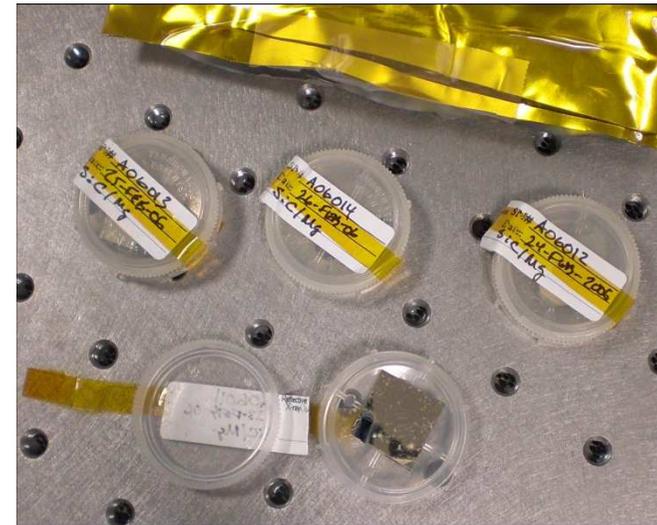
Multilayers for METIS

State of art

- **Mo/Si** At wavelengths longer than the Si – L absorption edge at 12,4 nm, Mo/Si multilayer are conventionally used for their high stability and fairly high reflectivity. However, reflectivity of the Mo/Si multilayer decreases gradually when the wavelength increases and the 30,4 nm reflectivity reduces to approximately 20%
- **Si/B₄C** Reflectance up to 25% can be obtained with **Si/B₄C** stack; however it has been verified to be unstable on the long term.
- **Mg/SiC** It could be a perfect candidate. It achieves up to 40% reflectance at 30,4 nm, and its stability has been deeply investigated. It has been selected to coat a mirror on board of the METIS prototype, the NASA Sounding – Rocket Coronagraphic Experiment (SCORE). ...It could be a perfect solution.... But....

Multilayers for METIS Mg/SiC ... why not?

- Mg/SiC samples with different thickness ratio between the two materials and capping layers characteristics have been fabricated .
- They offer superior performances in terms of reflectivity, but from our experience not always in terms of stability on the long run.
- For example, the regular handling and the atmosphere exposure affect the performances of the coatings.



SiC/Mg samples

In this scenario, we have investigated alternative solutions.

Multilayers for METIS

Alternative solutions

- ML with different capping layer (CL) materials (4 different types):

1. Mo/Si + CL 1
2. Mo/Si + CL 2
3. Mo/Si + CL 3
4. Mo/Si + CL 4

(Samples fabricated on Si wafer)

- Ir/Si ML (Samples fabricated on Si wafer)



New samples

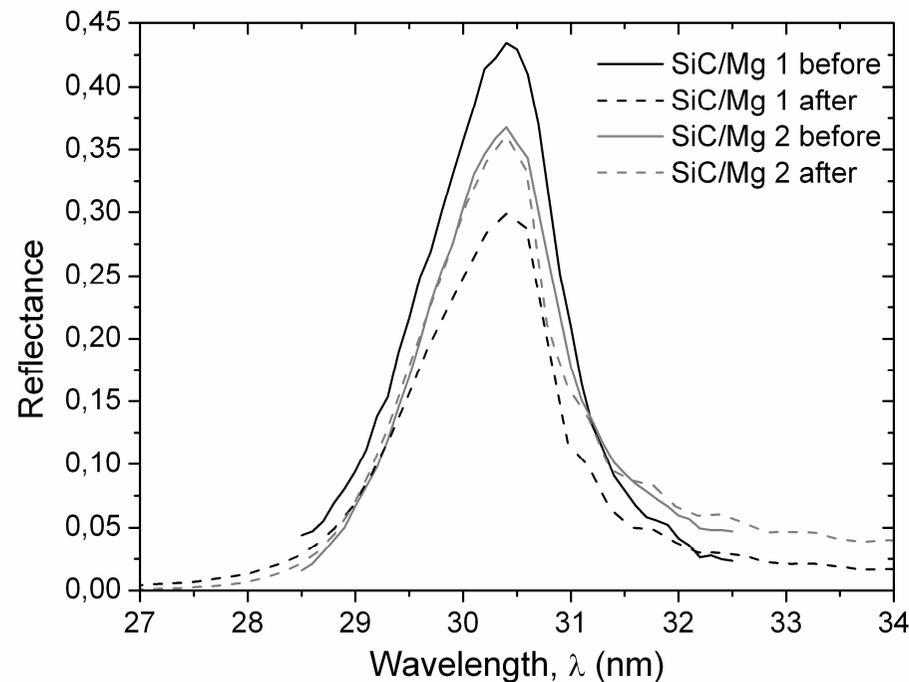
The samples have been designed at LUXOR Laboratory and deposited at Reflective X-ray Optics LLC (New York, USA) by DC magnetron sputtering onto polished Si (100) substrates*.

*Corso, A.J., Zuppella, P., Nicolosi, P., Windt, D.L., Gullikson, E., Pelizzo, M.G., "Capped multilayers with improved performances at 30,4 nm for future solar missions, Optics Express 2011

Multilayers for METIS

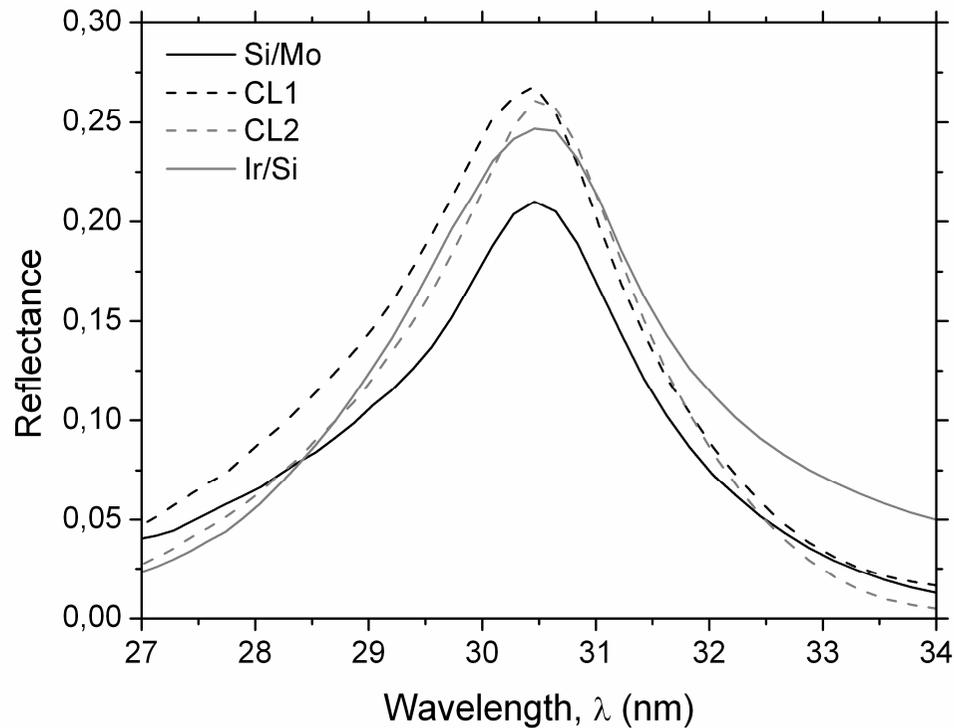
Measurements at 30,4 nm – Mg/SiC

The EUV measurements have been performed at BEAR beam line ELETTRA, Synchrotron - Trieste (Italy). Reflectance was measured at 5° from normal over the spectral range 27 – 34 nm. The figure refers to the reflectance of two different Mg/SiC samples after deposition and five years later.



Multilayers for METIS

Measurements at 30,4 nm – new samples



The reflectivity is stable over six months!

Multilayers for METIS

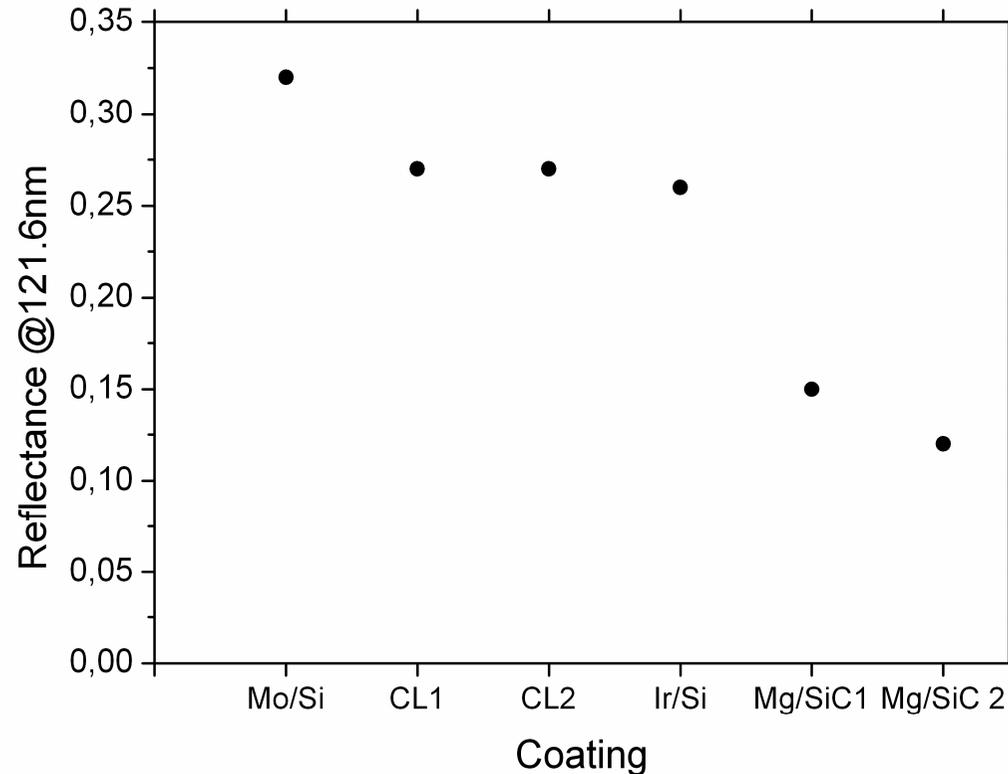
Measurements at 30,4 nm – considerations

- The new multilayers proposed perform better than the classical Mo/Si. The classical Mo/Si reflects 20% as expected , while the new structures proposed show better performances: 24% for Ir/Si multilayer and almost 27% for capping layers onto Mo/Si structures.
- After five years one of the Mg/SiC multilayers reflects 35%. However we can't fully explain the behavior of the Mg/SiC and foresee the total degradation of the reflectivity by aging.

Then, the new multilayers can be considered valid candidates for their stability and significant peak reflectance over time. **Further tests are in progress !**

Multilayers for METIS Measurements at 121,6 nm

The reflectance at 121,6 nm has been measured by using the normal incidence reflectometer at CNR – LUXOR (Padova, Italy)



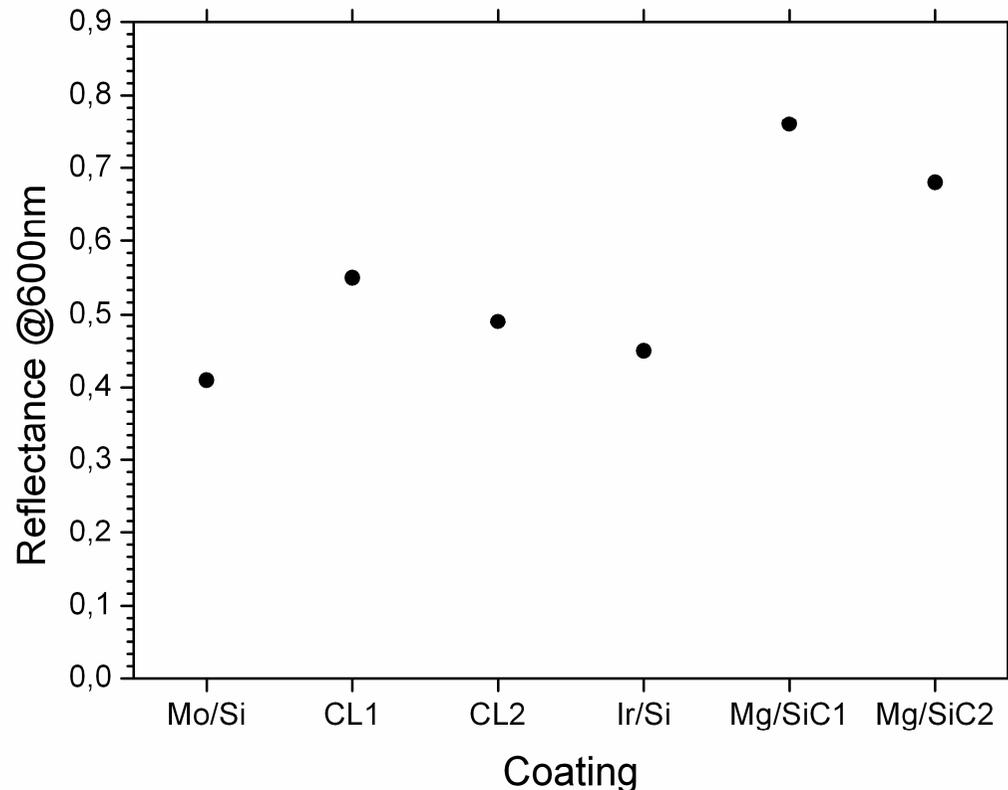
Notes:

- The Mo/Si performs better than the others multilayers: reflectivity of 32% has been measured, on the contrary no better than 15% for the Mg/SiC coatings

- After five years the reflectivity of the Mg/SiC decreased of 30%. The reflectance at 121,6 nm depends on the first layer of the ML. We will expect further reflectance loss for normal handling operations.

Multilayers for METIS Measurements in the VIS range

The VIS reflectance was carried on by using the Cary 5000 spectrophotometer in 250 – 800 nm spectral range. The figure below shows the experimental results at 600 nm.



Notes:

- The new samples are stable over time
- By looking on the graph , the Mg/SiC seems to give superior performances, but the measurements have been done before the normal handling operations.

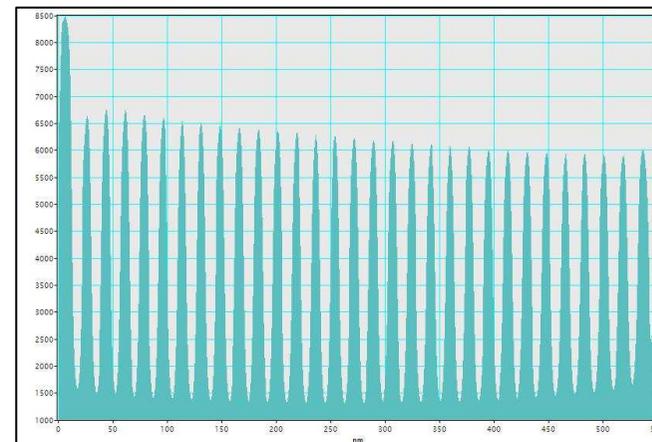
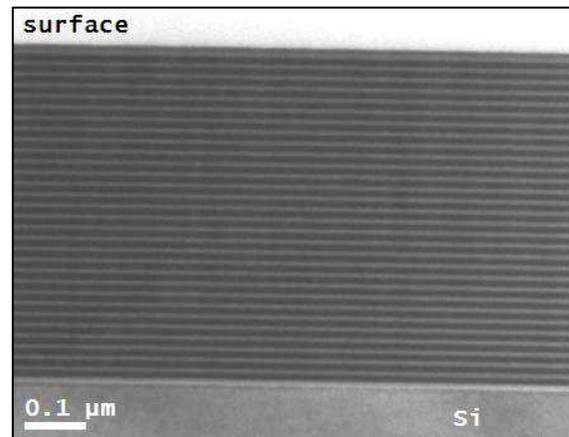
Multilayers for METIS

Structural analysis – Ir/Si

The structure and morphology of the innovative periodic Ir/Si multilayer have been investigated *.

Microscope:

FEI – Technai Field – emission Gun (FEG) TEM F – 20 Super Twin, operating at 200 KeV equipped with an EDX energy – dispersive X- ray spectrometer, and with scanning deflection coils that allow to form images in the scanning mode of operation (STEM)

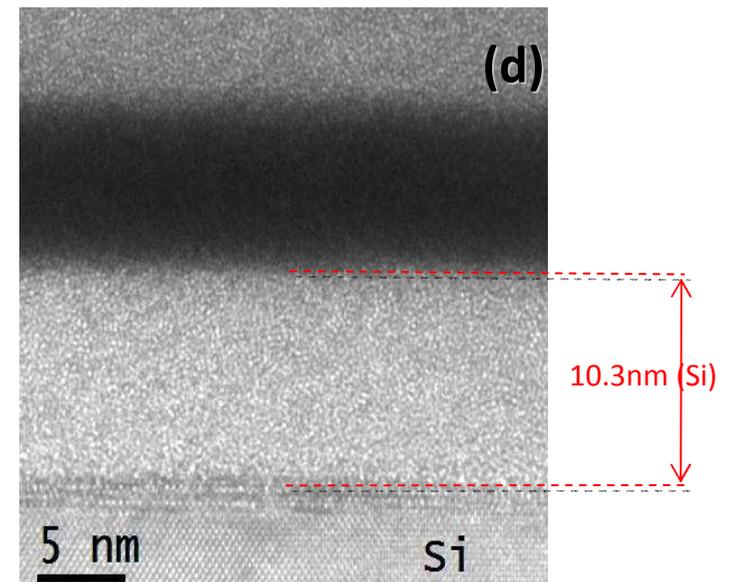
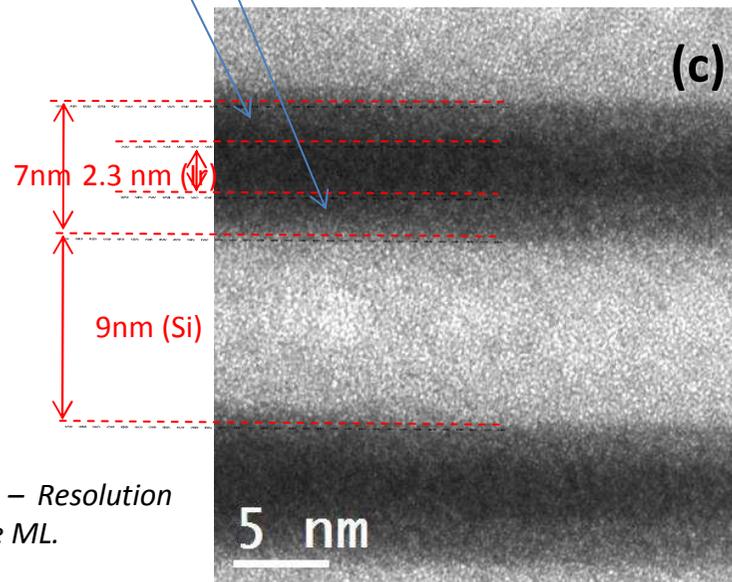
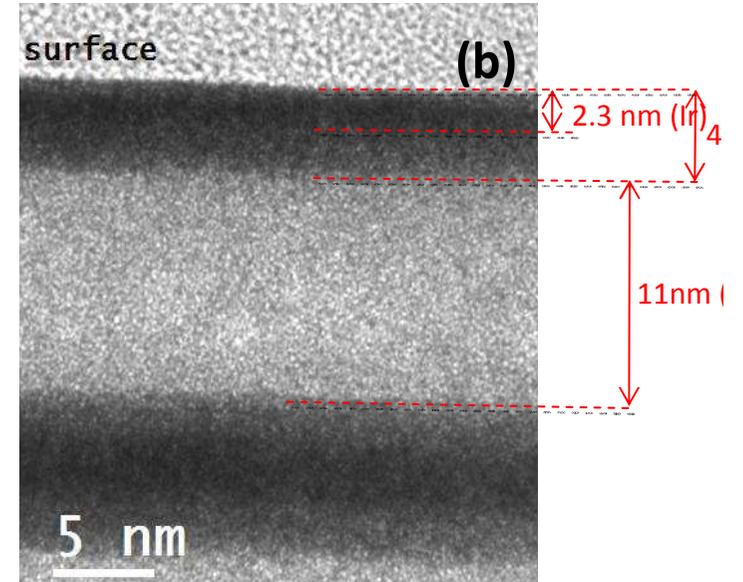
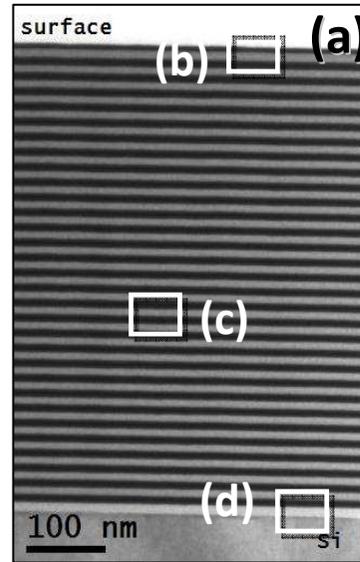


Bright-Field cross-sectional image of the Multilayer and the corresponding intensity line scan.
*Zuppella, P., Monaco, G., Corso, A.J., Nicolosi, P., Windt, D.L., Bello, V., Matter, G., Pelizzo, M.G., “Iridium /Silicon multilayers for EUV applications in the 20 – 35 nm wavelength range”, Opt. Lett. 36(7), 1203 – 1205 (2011)

Multilayers for METIS

Structural analysis – Ir/Si

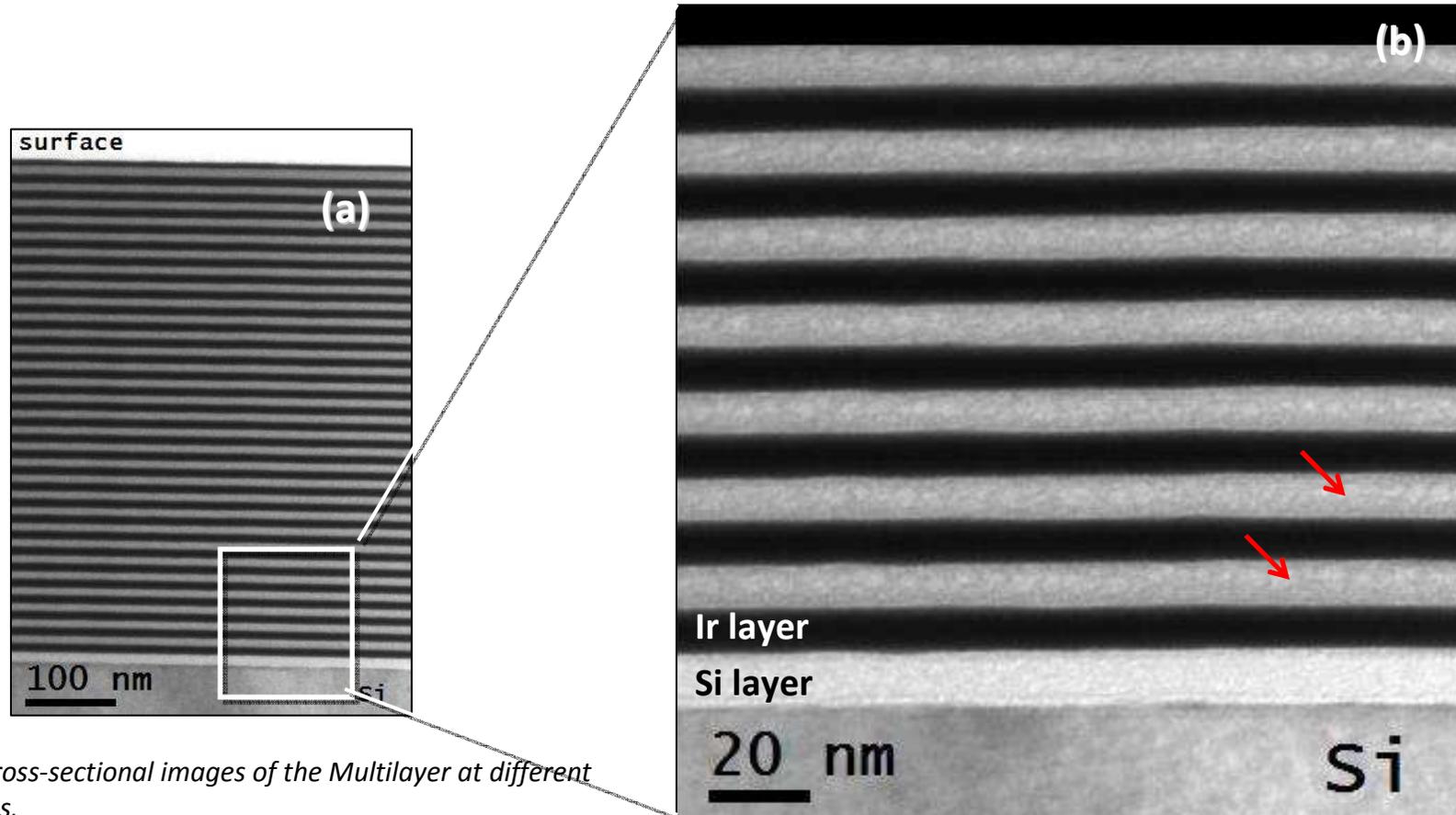
A mixed Si – Ir layer is present at each interface, the thickness is $t_{Si_Ir} = 2,4 \pm 0,4 \text{ nm}$



Bright – Field (a) and High – Resolution (b), (c), and (d) images of the ML.

Multilayers for METIS

Structural analysis – Ir/Si

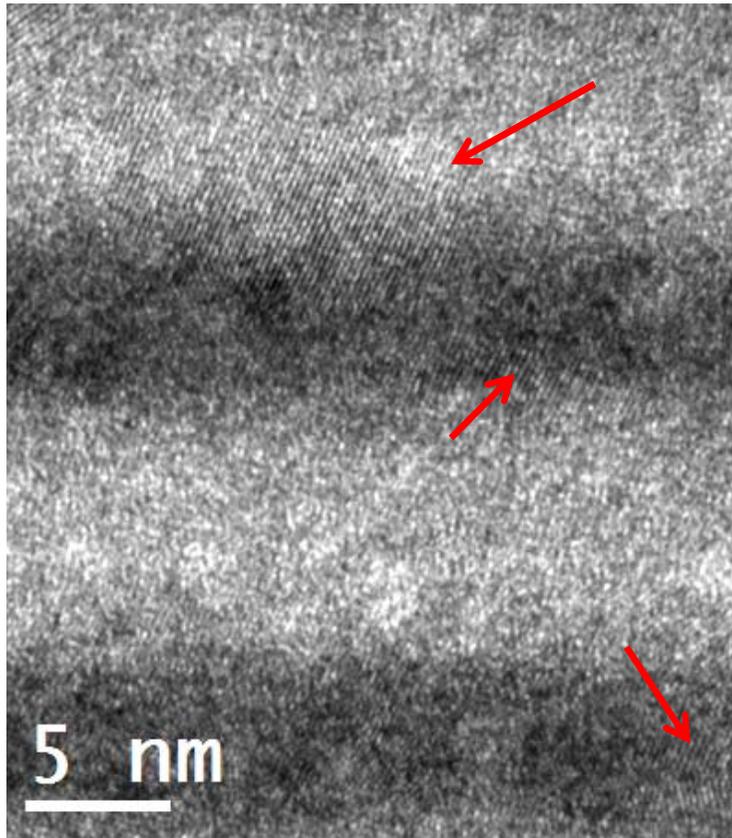


Bright-Field cross-sectional images of the Multilayer at different magnifications.

It is possible to observe modulations in the contrast within the a-Si layers, probably indicating that the a-Si layers are not perfectly densified (red arrows).

Multilayers for METIS

Structural analysis – Ir/Si



High-Resolution image acquired on a very thin region of the multilayer. Red arrows indicate Ir crystalline domains.

Experim. d(Å)	Ir-FCC d(Å)	Ir ₃ Si-Tetrag.d(Å)
2.17±0.05	2.22 (111)	2.18 (202)
1.91±0.03	1.92 (200)	

Tab.1 Experimental interplanar distances found by FFT analysis of the HR-images compared with that of Ir-fcc and Ir₃Si Tetragonal phase, with the corresponding (hkl) indices.

The HR-TEM image has been acquired on a thinner region: the red arrows indicate the presence of some crystalline domains with interplanar distances (reported in Tab.1) that could be attributed to an fcc-Ir phase. The last column of the tables show that an interplanar distance found by FFT analysis of HR images could be also attributed to the presence of a crystalline Ir₃Si phase.

Conclusions

- Multilayer design and development at LUXOR relative to ML's for Space Applications have been presented

ACKNOWLEDGEMENTS

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THANK YOU